### Clock dependence and unitarity in quantum cosmology

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# Work in collaboration with Steffen Gielen arXiv: 2005.05357 + 2109.02660

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- Quantisation(s) and unitarity
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# What is the problem of time (POT)?

Diffeomorphism invariance  $\implies$  GR is a constrained system:

$$\mathcal{S}_{EH} = \int \mathrm{d}t \mathrm{d}^3 x \; \left[ p^{ab} \dot{h}_{ab} - N \mathcal{H}^g_\perp - N^a \mathcal{H}^g_a 
ight]$$

N and  $N^a$  (lapse and shift functions) are Lagrange multipliers. Hence,

$$\mathcal{H}^{g}_{\perp}=0, \quad \mathcal{H}^{g}_{a}=0$$

When we quantise we find

$$\hat{\mathcal{H}}_{\perp}^{g}\Psi=0,\quad\hat{\mathcal{H}}_{a}^{g}\Psi=0$$

How do we make observables evolve?



# Possible approaches to the POT

#### Choose time before quantisation

Reduced quantisation

- Solve the constraint classically and quantise only the true degrees of freedom
- Ambiguities in gauge fixing result in different quantum theories<sup>1</sup>

#### Choose time after quantisation

We start from a WdW equation<sup>1</sup> and then:

- A Choose a clock<sup>2</sup> first and then an inner product
- B Choose a too big Hilbert space and reduce it<sup>3</sup> (Dirac quantisation)

#### All these approaches are theory independent

<sup>1</sup>Ordering and other issues <sup>2</sup>Why one and not another? <sup>3</sup>Steps not straight forward, e.g., group averaging



Clock dependence and unitarity in QC



- O Choose a cosmological model (⇐ easier metric)
- Quantise using the Wheeler–DeWitt equation
- S Choose several dynamical variables as clock
- Ompare the resulting theories
- Sompare with the Dirac quantisation
- Study singularity resolution



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## **Classical solutions**

Ingredients

• Flat FLRW: 
$$ds^2 = -N(\tau)^2 d\tau^2 + a(\tau)^2 (dx^2 + dy^2 + dz^2)$$

• Free massless scalar field  $\phi$ 

• A cosmological constant  $\Lambda$  from unimodular gravity

Unimodular gravity= GR+fixed determinant. Additional fields  $T^a$  can be introduced to recover full diffeomorphism invariance.

$$\begin{split} \mathcal{S}_{PUM} &= \int \mathrm{d}^4 x \left\{ \frac{\sqrt{-g}}{2\kappa} [R - 2\Lambda] + \Lambda \partial_a T^a - \frac{1}{2} g^{ab} \partial_a \phi \partial_b \phi \right\} \\ &= V_0 \int_{\mathbb{R}} \mathrm{d}\tau \left\{ \frac{3\dot{a}^2 a}{N\kappa} - N a^3 \frac{\Lambda}{\kappa} + \Lambda \dot{T} + \frac{a^3}{2N} \dot{\phi}^2 \right\} \end{split}$$



### Classical solutions continued

Change of variables:

$$\begin{aligned} \mathbf{v} &= 2\sqrt{\frac{V_0}{3}}\mathbf{a}^3, \ \pi_{\mathbf{v}} &= \sqrt{\frac{1}{12V_0}}\frac{\pi_{\mathbf{a}}}{\mathbf{a}^2}, \ \lambda &= V_0\Lambda, \ t = \frac{T}{V_0}, \\ \varphi &= 2\sqrt{\frac{3}{8}}\phi, \ \pi_{\varphi} &= \frac{1}{2}\sqrt{\frac{8}{3}}\pi_{\phi} \end{aligned}$$

where  $\{t, \lambda\} = 1$ . The cosmological constant is a constant of motion.

The action can be brought into Hamiltonian form with:

$$\mathcal{H} = \tilde{N} \left[ -\pi_v^2 + \frac{\pi_\varphi^2}{v^2} + \lambda \right] \stackrel{\text{constraint}}{\Longrightarrow} \mathcal{C} = -\pi_v^2 + \frac{\pi_\varphi^2}{v^2} + \lambda = 0$$

Choice:  $\tilde{N} = 1 \implies \frac{\mathrm{d}t}{\mathrm{d}\tau} = 1$ . *v* volume



#### The three different clocks

For  $\lambda > 0$  the classical solutions are

$$\begin{split} v(t) &= \sqrt{-\frac{\pi_{\varphi}^2}{\lambda} + 4\lambda(t - t_0)^2}, \qquad \varphi(t) = \frac{1}{2} \log \left| \frac{\pi_{\varphi} - 2\lambda(t - t_0)}{\pi_{\varphi} + 2\lambda(t - t_0)} \right| + \varphi_0 \\ v(\varphi) &= \frac{|\pi_{\varphi}|}{\sqrt{\lambda}|\sinh(\varphi - \varphi_0)|}, \qquad t(\varphi) = -\frac{\pi_{\varphi}}{2\lambda} \coth(\varphi - \varphi_0) + t_0 \\ t(v) &= t_0 - \operatorname{sgn}(\pi_v) \frac{1}{2} \sqrt{\frac{v^2}{\lambda} + \frac{\pi_{\varphi}^2}{\lambda^2}}, \quad \varphi(v) = \varphi_0 + \log \left| \frac{\pi}{\sqrt{\lambda}v} + \sqrt{\frac{\pi_{\varphi}^2}{\lambda v^2} + 1} \right| \end{split}$$

The big bang/big crunch singularity is at  $t_{sing} = \frac{|\pi_{\varphi}|}{2\lambda}$ ,  $\log \frac{v_{sing}}{v_0} = -\infty$ ,  $\varphi_{sing} = \pm \infty$ Spatial infinity is at  $t_{\infty} = \pm \infty$ ,  $\log \frac{v_{\infty}}{v_0} = \infty$ ,  $\varphi_{\infty} = \varphi_0$ 



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#### Definition (Slow (fast) clocks)

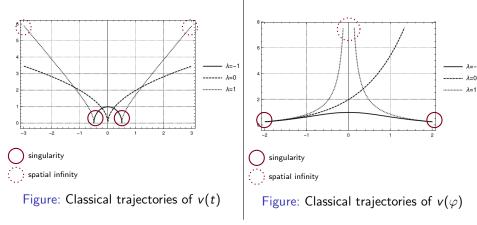
Let  $x \in \mathbb{R}$  be a dynamical variable used to express the variation of the remaining parameters of the universe. x is said to be *slow (or fast)* at the singularity/spatial infinity if the singularity/spatial infinity is reached at a *finite value*  $x_0$  (or at  $\pm \infty$ )

In other words: If a clock is able to "push" the singularity/spatial infinity to the boundaries of its domain it is fast

- t-clock: slow at the singularity and fast at spatial infinity
- $\varphi$ -clock: fast at the singularity and slow at spatial infinity
- *v*-clock: fast everywhere



# Trajectories of v(t) and $v(\varphi)$





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 $\lambda = -1$ 

-· \lambda\_0

Gotay and Demaret (1983) Conjecture:

- Classically: slow clock  $\implies$  dynamics are truncated
- When quantising and demanding unitarity: state norm must be well defined everywhere in clock space ⇒ artificially extending the solution to the whole domain of the clock ⇒ singularity (or spatial infinity) resolution

We are going to verify this conjecture in our work



- Pawłowski and Ashtekar (2012) arXiv: 1011.3022  $\rightarrow$  Similar model with a fixed cosmological constant form a  $\varphi$ -clock perspective
- **Gryb** and Thébault (2019) arXiv: 1801.05789 and 1801.05826  $\rightarrow$  Same model from a *t*-clock perspective
- Gielen and Turok (2016) arXiv: 1510.00699  $\rightarrow$  Same model with a *v*-clock perspective
- Bojowald, Brizuela, Hernandez, Koop and Morales-Tecolt (2011) arXiv: 1011.3022 → Similar model with fixed cosmological constant using momenta



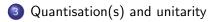
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### Wheeler–DeWitt equation

$$\mathcal{C} = -\pi_v^2 + \frac{\pi_\varphi^2}{v^2} + \lambda = 0 \implies g^{AB}\pi_A\pi_B + \lambda = 0, \quad g^{AB} = \begin{pmatrix} -1 & 0 \\ 0 & \frac{1}{v^2} \end{pmatrix}$$

Ordering problem: how do we go from here to a partial differential equation?

Answer: Replace  $g^{AB}\pi_A\pi_B$  (which is flat) by the Laplace Beltrami operator  $-\hbar^2\Box$ 

$$\hat{\mathcal{C}}\Psi = 0 \implies \left(\hbar^2 \frac{\partial^2}{\partial v^2} + \frac{\hbar^2}{v} \frac{\partial}{\partial v} - \frac{\hbar^2}{v^2} \frac{\partial^2}{\partial \varphi^2} - i\hbar \frac{\partial}{\partial t}\right) \Psi(v, \varphi, t) = 0$$
(WdW)

POT: how to extract time evolution from this equation? Answer: A and B (and more?)



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### Answer A: choose a clock

 $t,\,\varphi$  and v are good clocks classically, hence both can a priori be good clocks for the quantum theory

The *t*-clock theory  

$$i\hbar\partial_t \Psi = -\underbrace{\left(\hbar^2 \partial_v^2 + \frac{\hbar^2}{v} \partial_v - \frac{\hbar^2}{v^2} \partial_\varphi^2\right)}_{\hat{\mathcal{H}}} \Psi$$
Multiply (WdW) by  $v^2$   

$$\hbar^2 \partial_\varphi^2 \Psi = \underbrace{\left(\hbar^2 (v \partial_v)^2 - i v^2 \hbar \partial_t\right)}_{\hat{\mathcal{G}}} \Psi$$
K.G. eq.  $\Longrightarrow$  K.G. inner product  

$$\langle \Psi | \Phi \rangle_t = \int d\varphi dv \ v \bar{\Psi} \Phi$$

$$\langle \Psi | \Phi \rangle_\varphi = i \int dt \frac{dv}{v} \left(\bar{\Psi} \partial_\varphi \Phi - \Phi \partial_\varphi \bar{\Psi}\right)$$

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the v-clock theory

Multiply (WdW) by  $v^2$ 

$$\hbar^{2}(\nu\partial_{\nu})^{2}\Psi = \underbrace{\left(\hbar^{2}\partial_{\varphi}^{2} + i\nu^{2}\hbar\partial_{t}\right)}_{\hat{\mathcal{F}}}\Psi$$

K.G. eq.  $\implies$  K.G. inner product

$$\langle \Psi | \Phi \rangle_{\mathbf{v}} = i \int \mathrm{d}t \mathrm{d}\varphi \,\, \mathbf{v} \left( \bar{\Psi} \partial_{\mathbf{v}} \Phi - \Phi \partial_{\mathbf{v}} \bar{\Psi} \right)$$



### Choose a clock (continued)

The solutions to (WdW) are Bessel functions

$$\begin{split} \Psi(\mathbf{v},\varphi,t) &= \int \frac{\mathrm{d}\lambda}{2\pi\hbar} \frac{\mathrm{d}k}{2\pi} e^{ik\varphi} e^{i\lambda\frac{t}{\hbar}} \left[ \alpha(k,\lambda) J_{i|k|} \left( \frac{\sqrt{\lambda}}{\hbar} \mathbf{v} \right) + \beta(k,\lambda) J_{-i|k|} \left( \frac{\sqrt{\lambda}}{\hbar} \mathbf{v} \right) \right] \\ &+ \int \frac{\mathrm{d}\lambda}{2\pi\hbar} \frac{\mathrm{d}\kappa}{2\pi} e^{\kappa\varphi} e^{i\lambda\frac{t}{\hbar}} \left[ \gamma(\kappa,\lambda) J_{|\kappa|} \left( \frac{\sqrt{\lambda}}{\hbar} \mathbf{v} \right) + \epsilon(\kappa,\lambda) J_{-|\kappa|} \left( \frac{\sqrt{\lambda}}{\hbar} \mathbf{v} \right) \right] \end{split}$$

 $\alpha(k,\lambda)$ ,  $\beta(k,\lambda)$ ,  $\gamma(\kappa,\lambda)$  and  $\epsilon(\kappa,\lambda)$  need to be constrained for two main reasons

- Normalisation:  $\langle \Psi | \Phi \rangle < \infty$
- Time independence: We want our theory to be unitary



#### Normalisation:

$$\begin{array}{c|c} t\text{-clock} & \varphi\text{-clock}^* & v\text{-clock}^* \\ \gamma(\kappa,\lambda) = \epsilon(\kappa,\lambda) = 0 & \epsilon(\kappa,\lambda) = 0 \end{array}$$

\*Klein–Gordon inner products are not positive definite but can easily be changed as solutions are already separating positive and negarive norm contributions



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# Choose a clock (continued)

#### Time independence

$$\begin{split} \partial_{t} \langle \Psi | \Phi \rangle_{t} &= 0 \iff \left\langle \hat{\mathcal{H}} \Psi \middle| \Phi \right\rangle_{(L^{2}, v d v d \varphi)} = \left\langle \Psi \middle| \hat{\mathcal{H}} \Phi \right\rangle_{(L^{2}, v d v d \varphi)} \\ \partial_{\varphi} \langle \Psi | \Phi \rangle_{\varphi} &= 0 \iff \left\langle \hat{\mathcal{G}} \Psi \middle| \Phi \right\rangle_{(L^{2}, \frac{d v}{v} d t)} = \left\langle \Psi \middle| \hat{\mathcal{G}} \Phi \right\rangle_{(L^{2}, \frac{d v}{v} d t)} \\ \partial_{v} \langle \Psi | \Phi \rangle_{v} &= 0 \iff \left\langle \hat{\mathcal{F}} \Psi \middle| \Phi \right\rangle_{(L^{2}, v d t d \varphi)} = \left\langle \Psi \middle| \hat{\mathcal{F}} \Phi \right\rangle_{(L^{2}, v d t d \varphi)} \end{split}$$

In a nutshell,

the unitarity requirement is equivalent to a self-adjointness problem



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Symmetric operators fall in three categories:

- Self-adjoint operators (like  $\hat{\mathcal{F}}$ )
- Are not self-adjoint but admit self-adjoint extensions (like  $\hat{\mathcal{G}}$  or  $\hat{\mathcal{H}}$ )
- Are not self-adjoint and do not admit self-adjoint extensions

Because  $\hat{\mathcal{F}}$  is already self-adjoint, we are done for the v-clock theory



# Choose a clock (continued)

 $\hat{\mathcal{H}}$  and  $\hat{\mathcal{G}}$  are not self-adjoint  $\implies$  boundary conditions have to be imposed *t*-clock theory  $\varphi$ -clock theory  $\int \mathrm{d}\varphi \left[ v \bar{\Psi} \partial_v \Phi - v \Phi \partial_v \bar{\Psi} \right]_{v=0} = 0$  $\int \mathrm{d}t \left[ v \bar{\Psi} \partial_v \Phi - v \Phi \partial_v \bar{\Psi} \right]^{v=\infty} = 0$ Real combinations of Bessel Real combination of Bessel functions functions  $\Psi_{t} \sim \int \frac{\mathrm{d}\lambda}{2\pi\hbar} \frac{\mathrm{d}k}{2\pi} e^{ik\varphi} e^{i\lambda\frac{t}{\hbar}} \alpha(k,\lambda) \operatorname{Re} \left[ e^{i\frac{\vartheta(k)-i|k|\log\sqrt{\frac{\lambda}{\lambda_{0}}}}} J_{i|k|}\left(\frac{\sqrt{\lambda}}{\hbar}v\right) \right]$  $\Psi_{\varphi} \sim \int \frac{\mathrm{d}\lambda}{2\pi\hbar} \frac{\mathrm{d}k}{2\pi} e^{ik\varphi} e^{i\lambda\frac{t}{\hbar}} \alpha(k,\lambda) \operatorname{Re} \left[ \sqrt{\frac{\sinh\left((|k| - i\kappa_0(\lambda))\frac{\pi}{2}\right)}{\sinh\left((|k| + i\kappa_0(\lambda))\frac{\pi}{2}\right)}} J_{i|k|}\left(\frac{\sqrt{\lambda}}{\hbar}v\right) \right]$ 

# Choose a clock (conclusion)

Remark:

- In the *t*-theory, the boundary condition is at v = 0 (classical singuarity), where classically the clock is slow
- In the  $\varphi$ -theory, the boundary condition is at  $v = \infty$  (spatial infinity), where again classically the clock is slow

In conclusion:

- Unitarity implies a boundary condition at the slow point of the clock
- O Different clock choices lead to different boundary conditions

Is this a consequence of choosing a clock first and then building an inner product?

#### What about Dirac quantisation?



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### Answer B: Dirac quantisation

We have two WdW equations

$$\hat{\mathcal{C}}_{1}\Psi = 0, \implies \left(\hbar^{2}\frac{\partial^{2}}{\partial v^{2}} + \frac{\hbar^{2}}{v}\frac{\partial}{\partial v} - \frac{\hbar^{2}}{v^{2}}\frac{\partial^{2}}{\partial \varphi^{2}} - i\hbar\frac{\partial}{\partial t}\right)\Psi = 0 \quad (WdW1)$$
$$\hat{\mathcal{C}}_{2}\Psi = 0, \implies \left(\hbar^{2}\left(v\frac{\partial}{\partial v}\right)^{2} - \hbar^{2}\frac{\partial^{2}}{\partial \varphi^{2}} - i\hbar v^{2}\frac{\partial}{\partial t}\right)\Psi = 0 \quad (WdW2)$$

Note that (WdW2)=  $v^2$ (WdW1) The interpretation of the second order derivatives as a  $\hbar^2 \Box$  for different metrics motivates different kinematical inner products

$$\langle \Psi | \Phi \rangle_{\mathsf{kin}_{1}} = \int \mathrm{d}t \mathrm{d}\varphi \mathrm{d}v \ v \bar{\Psi} \Phi$$
 (IP1)  
 
$$\langle \Psi | \Phi \rangle_{\mathsf{kin}_{2}} = \int \mathrm{d}t \mathrm{d}\varphi \frac{\mathrm{d}v}{v} \ \bar{\Psi} \Phi$$
 (IP2)

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# Dirac quantisation (continued)

Demanding  $\hat{C}_1$  and  $\hat{C}_2$  to be self-adjoint with respect to the inner products (IP1) and (IP2) corresponds to self-adjointness of the one dimensional operators

$$\hat{\mathcal{D}}_1 = \hbar^2 \left( -\frac{\partial^2}{\partial v^2} - \frac{k^2 + \frac{1}{4}}{v^2} \right), \quad \hat{\mathcal{D}}_2 = -\hbar^2 \left( v \frac{\partial}{\partial v} \right)^2 - \lambda v^2$$

Recall,

$$\hat{\mathcal{H}} = \left(\hbar^2 \frac{\partial^2}{\partial v^2} + \frac{\hbar^2}{v} \frac{\partial}{\partial v} - \frac{\hbar^2}{v^2} \frac{\partial^2}{\partial \varphi^2}\right), \quad \hat{\mathcal{G}} = \left(\hbar^2 \left(v \frac{\partial}{\partial v}\right)^2 - i v^2 \hbar \frac{\partial}{\partial t}\right)$$

Despite being two dimensional, study  $\hat{\mathcal{H}}$  and  $\hat{\mathcal{G}}$  reduces to analysing  $\hat{\mathcal{D}}_1$  and  $\hat{\mathcal{D}}_2$ 



#### The self-adjointness problem remains in the Dirac quantisation

We multiplied the Wheeler-DeWitt equation by a phase function to find different theories, but  $\hat{C}\Psi = 0$  and  $\hat{N}\hat{C}\Psi = 0$  have the same solutions, this is consequence of reparametrisation invariance

How do we choose between the different Wheeler-DeWitt equations?



■ P. A. Höhn, A. R. H. Smith and M. P. Lock 2019 and 2021 arXiv:1912.00033 and 2007.00580  $\rightarrow$  framework for clock changes

Our work *does not* contradict theirs as the framework is different: they work with clock changes given a WdW equation and instead we develop a theory for two different WdW equations



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### Dynamics of the v-theory

General solution:

$$\Psi(v,\varphi,t)_{sc} = \int \frac{\mathrm{d}\lambda}{2\pi\hbar} \frac{\mathrm{d}k}{2\pi} e^{ik\varphi} e^{i\lambda\frac{t}{\hbar}} \left[ \alpha(k,\lambda) J_{i|k|} \left( \frac{\sqrt{\lambda}}{\hbar} v \right) + \beta(k,\lambda) J_{-i|k|} \left( \frac{\sqrt{\lambda}}{\hbar} v \right) \right]$$

- $\alpha(k,\lambda) = 0$  (only outgoing modes)
- $\beta(k,\lambda)$  sharply peaked around  $k = k_c$  and  $\lambda = \lambda_c$

Expanding around v = 0

- Classically:  $t(v) = \frac{\hbar |k_c|}{2\lambda_c} + \frac{v^2}{4\hbar |k_c|} \frac{\lambda_c v^4}{16\hbar^3 |k_c|^3} + O(v^6)$
- Quantum theory:  $\langle \Psi_{sc}(v) | t | \Psi_{sc}(v) \rangle = \langle t(v) \rangle_{\Psi_{sc}} = \frac{\hbar |k_c|}{2\lambda_c} + \frac{v^2}{4\hbar |k_c|} - \frac{\lambda_c v^4}{16\hbar^3 (|k_c| + |k_c|^3)} + O(v^6)$

#### No strong divergence from the classical theory



Recall the general solution

$$\Psi_{sc} \sim \int \frac{\mathrm{d}\lambda}{2\pi\hbar} \frac{\mathrm{d}k}{2\pi} \ e^{ik\varphi} e^{i\lambda\frac{t}{\hbar}} \alpha(k,\lambda) \operatorname{Re}\left[ e^{i\frac{\vartheta(k)}{\hbar} - i|k|\log\sqrt{\frac{\lambda}{\lambda_0}}} J_{i|k|}\left(\frac{\sqrt{\lambda}}{\hbar}v\right) \right]$$

- $\vartheta(k) = 0$  for simplicity
- $\alpha(k,\lambda)$  sharply peaked around  $k=k_c$  and  $\lambda=\lambda_c$
- expectation values =  $\langle \Psi_{sc}(t) | \, v \, | \Psi_{sc}(t) 
  angle = \langle v(t) 
  angle_{\psi_{sc}}$
- criteria for singularity resolution:  $\langle v(t) 
  angle_{\psi_{sc}} > C > 0$



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Bessel function asymptote for small arguments

$$J_{i|k|}\left(\frac{\sqrt{\lambda}}{\hbar}\nu\right) \longrightarrow \frac{e^{i|k|\log\left(\frac{\sqrt{\lambda}}{2\hbar}\nu\right)}}{\Gamma(1+i|k|)}$$

With the prefactor

$$e^{i\vartheta(k)-i|k|\log\sqrt{rac{\lambda}{\lambda_0}}}$$

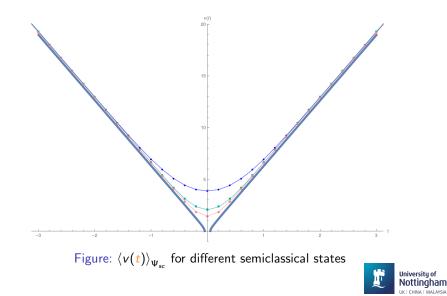
 $\Psi_{sc}$  looks like a superposition of plane waves independent of  $\lambda$  outgoing from and ingoing to the singularity



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### Singularity resolution



### Dynamics of the $\varphi$ -clock theory

Recall the general solution

$$\Psi_{sc} \sim \int \frac{\mathrm{d}\lambda}{2\pi\hbar} \frac{\mathrm{d}k}{2\pi} e^{ik\varphi} e^{i\lambda\frac{t}{\hbar}} \alpha(k,\lambda) \operatorname{Re}\left[\sqrt{\frac{\sinh\left(\left(|k|-i\kappa_{0}(\lambda)\right)\frac{\pi}{2}\right)}{\sinh\left(\left(|k|+i\kappa_{0}(\lambda)\right)\frac{\pi}{2}\right)}} J_{i|k|}\left(\frac{\sqrt{\lambda}}{\hbar}\nu\right)\right]$$

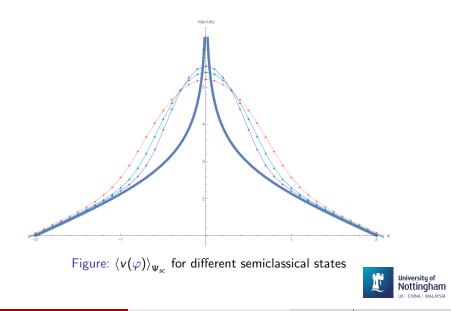
- $\kappa_0(\lambda) = 0$  for simplicity
- $\alpha(k,\lambda)$  sharply peaked around  $k = k_c$  and  $\lambda = \lambda_c$
- expectation values  $\langle \Psi_{sc}(\varphi) | v | \Psi_{sc}(\varphi) \rangle = \langle v(\varphi) \rangle_{\Psi_{sc}}$  and  $\langle \Psi_{sc}(\varphi) | t | \Psi_{sc}(\varphi) \rangle = \langle t(\varphi) \rangle_{\Psi_{sc}}$
- criteria for quantum recollapse  $\langle 
  u(arphi) 
  angle_{\psi_{sc}} < \mathcal{C}' < \infty$



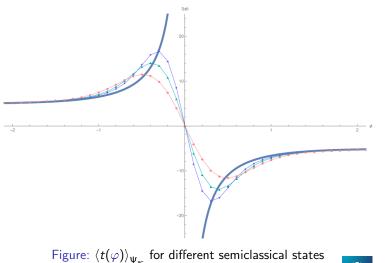
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### Quantum recollapse



### Quantum recollapse





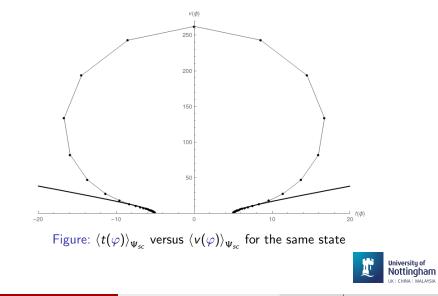
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### Quantum recollapse



- The v-clock shows no strong divergence from the classical theory
- The *t*-clock theory shows singularity resolution  $(\langle v(t) \rangle_{\Psi_{sc}} > C > 0)$
- The  $\varphi$ -clock theory shows a quantum recollapse  $(\langle v(\varphi) \rangle_{\Psi_{sc}} < C' < \infty)$
- The quantum behaviour can be interpreted as a reflection from the boundary (either v = ∞ or v = 0) ⇒ The boundary condition "decides" whether the theory is singular or not
- Far from the boundary, the classical and quantum curves look very similar



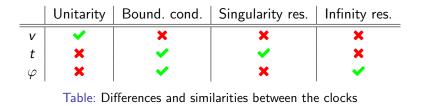
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### A few things to take from this talk

- The problem of time has many nuances
- Quantum cosmology is a good testing ground
- Unitarity requirements are what lead to different theories and,
- They are present in different quantisation schemes
- Singularity resolution is not a feature of the theory, it is a feature of the clock ⇒ clock choices should be given more attention when studying these models



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# Thank you!

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