

# The imaginary part of the GR action and the large-spin 4-simplex amplitude

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# Summary

## Classical GR

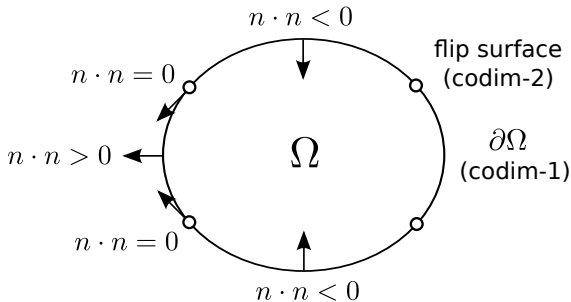
- First-order and second-order actions agree on-shell.
- The action has a nonvanishing imaginary part:

$$\text{Im } S = \frac{1}{16G} \sum_{\text{flips}} A_{\text{flip}} = \frac{1}{4} \sum_{\text{flips}} \sigma_{\text{flip}}$$

## The large-spin limit of LQG

- The large-spin vertex amplitude reproduces the classical action iff analytically continued to  $\gamma = \pm i$ .
- A first-ever agreement in LQG between BH entropy derived from an effective action and from a microscopic state counting.
- The limit should be thought of as a high-energy (“transplanckian”) regime.

## Generic spacetime regions $\Omega$ with boundary $\partial\Omega$



- $\partial\Omega$  doesn't have to be at infinity!

### Normal vector $n^\mu$

- Sign of  $n^\mu$ : chosen so that  $n_\mu v^\mu > 0$  for outgoing  $v^\mu$ .
- Sign of  $n \cdot n$ : has to flip between spacelike and timelike patches.  
 $\Rightarrow n \cdot n$  cannot be constant everywhere!

# The second-order GR action

$$S = \frac{1}{16\pi G} \left( \int_{\Omega} \sqrt{-g} R d^4x + 2 \int_{\partial\Omega} \sqrt{\frac{-h}{n \cdot n}} K d^3x \right)$$

- The independent variable is  $g_{\mu\nu}$ . All the rest derived from it.
- Bulk term: Einstein-Hilbert.
- Boundary term: York-Gibbons-Hawking. [York, 1972; Gibbons & Hawking, 1977]
  - Defining property:  $\delta S = 0$  for  $\delta g_{\mu\nu}|_{\partial\Omega} = 0$  (no assumptions on  $\partial_{\mu}g_{\nu\rho}$ ).

## Boundary term notation

- $h_{ab}$  - intrinsic metric of boundary  $\partial\Omega$ ;  $h = \det h_{ab}$ .
- $n^{\mu}$  - boundary normal;  $K_a^b = \nabla_a n^b$  - extrinsic curvature;  $K = K_a^a$ .
- Valid for arbitrary norm and signature of  $n^{\mu}$ .
  - Rescaling  $n^{\mu}$  is a gauge freedom.

## Alternative - a densitized normal

For a clean variational principle, would be nicer to have a normal that is a unique function of  $\partial\Omega$  and  $g_{\mu\nu}$  everywhere.

- Can't use a unit normal: not well-defined at signature flips.

- Can use a densitized version:  $N_\mu = \frac{1}{3!} \sqrt{-g} \epsilon_{\mu abc} \epsilon^{abc}$ .

- Related to undensitized  $n^\mu$  as:  $N_\mu = \sqrt{\frac{-h}{n \cdot n}} n_\mu$ .

- However, covariant derivatives of  $N_\mu$  are awkward.

- Will be easier in the first-order formalism.

# The first-order Palatini action

$$S = \frac{1}{16\pi G} \left( \int_{\Omega} \Sigma^{IJ} \wedge F_{IJ} - \int_{\partial\Omega} \Sigma^{IJ} \wedge \left( A_{IJ} - \frac{2N_I dN_J}{N \cdot N} \right) \right)$$

## The independent variables

- Vielbein  $e^I_{\mu}$  and connection  $A^I_{\mu}$ .
- In form notation, just  $e^I$  and  $A^{IJ}$ .

## Useful derived quantities

- $\Sigma^{IJ} = \frac{1}{2} \epsilon^{IJKL} e_K \wedge e_L = \star(e \wedge e)^{IJ}$ .
- Curvature:  $F_{IJ} = dA_{IJ} + A_I^K \wedge A_{KJ}$ .
- Densitized normal:  $N_I = \frac{1}{3!} \epsilon^{abc} \epsilon_{IJKL} e_a^J e_b^K e_c^L$ .

## The first piece of the boundary term

$$S_{\partial\Omega, \text{old}} = -\frac{1}{16\pi G} \int_{\partial\Omega} \Sigma^{IJ} \wedge A_{IJ}$$

### Why have this term?

- Cancels the  $A_{IJ}$  variation from the bulk term.
- Thus, enforces the variational principle:  $\delta S = 0$  for  $\delta e^I|_{\partial\Omega} = 0$  (no assumptions on  $A_{IJ}$ ).

### Not gauge invariant

- Invariant if the direction of  $N_I$  is constant along the boundary.
- Cannot hold everywhere on a closed boundary, because of signature flips!

## The new, second piece of the boundary term

$$S_{\partial\Omega, \text{new}} = \frac{1}{8\pi G} \int_{\partial\Omega} \Sigma^{IJ} \wedge \frac{N_I dN_J}{N \cdot N} = \frac{1}{8\pi G} \int_{\partial\Omega} \Sigma^{IJ} \wedge \frac{n_I dn_J}{n \cdot n}$$

In second expression, introduced an undensitized normal  $n_I$  (free to rescale).

### Why have this term?

- For gauge invariance - cancels the gauge transformation of the  $\Sigma^{IJ} \wedge A_{IJ}$  term.

### Properties of the $N_I dN_J$ term

- Non-polynomial, due to the  $N \cdot N$  in the denominator.
- Depends only on the boundary value of  $e^I$ .  
 $\Rightarrow$  Doesn't affect the variational principle.



# On-shell equality of the Einstein-Hilbert and Palatini actions

## Bulk and boundary terms are equal separately

- $\sqrt{-g} R d^4x = \Sigma^{IJ} \wedge F_{IJ}$
- $\sqrt{\frac{-h}{n \cdot n}} K d^3x = -\frac{1}{2} \Sigma^{IJ} \wedge \left( A_{IJ} - \frac{2n_I dn_J}{n \cdot n} \right)$
- The boundary terms are equal thanks to the new  $n_I dn_J$  piece.

## Requires “half” of the equations of motion

- $T_I \equiv de_I + A_I{}^J e_J = 0.$   
(from variation of the first-order action w.r.t.  $A_{IJ}$ )

## The first-order action with Holst term

$$S = \frac{1}{16\pi G} \left( \int_{\Omega} \Sigma^{IJ} \wedge F_{IJ} - \int_{\partial\Omega} \Sigma^{IJ} \wedge \left( A_{IJ} - \frac{2N_I dN_J}{N \cdot N} \right) \right) + \frac{1}{16\pi G \gamma} \left( \int_{\Omega} e^I \wedge e^J \wedge F_{IJ} - \int_{\partial\Omega} e^I \wedge (e^J \wedge A_{IJ} - de_I) \right)$$

### The $e^I \wedge e^J \wedge A_{IJ}$ piece of the boundary term

- Cancels the  $A_{IJ}$  variation from the  $e^I \wedge e^J \wedge F_{IJ}$  bulk term.
- Thus, enforces the variational principle:  $\delta S = 0$  for  $\delta e^I|_{\partial\Omega} = 0$ .

### The $e^I \wedge de_I$ piece of the boundary term

- Combines with the  $e^I \wedge e^J \wedge A_{IJ}$  term to give  $e^I \wedge T_I$ .
- Thus, restores gauge invariance.
- Unlike the  $N_I dN_J$  piece, is polynomial.

# Holst action - properties

## Both bulk and boundary pieces of the Holst term vanish on-shell

- Again, assume “half” the equations of motion  $T^I = 0$ .
- Bulk piece then vanishes due to Bianchi identity  $e^J \wedge F_{IJ} = 0$ .
- Boundary piece vanishes trivially.

## The Holst term is not topological

- The boundary piece alone *is* topological.
  - $d(e^I \wedge T_I)$  is the Nieh-Yan density.
- However, only the correct bulk+boundary combination gives a good variational principle.

## Holst action - alternative expression

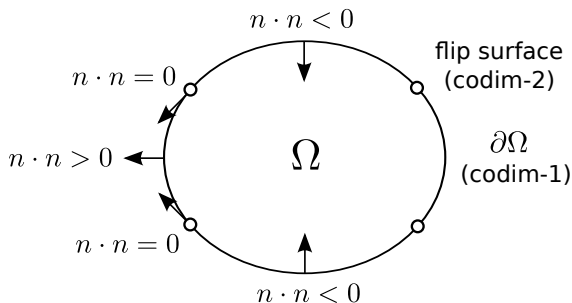
$$S = \frac{1}{16\pi G} \left( \int_{\Omega} \left( \star + \frac{1}{\gamma} \right) (e \wedge e)^{IJ} \wedge F_{IJ} - \int_{\partial\Omega} \left( \star + \frac{1}{\gamma} \right) (e \wedge e)^{IJ} \wedge (A_{IJ} - \Gamma_{IJ}^H) \right)$$

- $\Gamma_{IJ}^H$  is Peldan's hybrid connection (a non-polynomial functional of  $e^I$ ).  
[Peldan,1994]
- Manifest Hodge-duality structure, gauge invariance.
- Credit to Andreas Thurn.

## Summary on GR actions

- We now have the correct boundary terms for the Palatini and Holst actions.
- The correct Holst term vanishes on-shell, but is not topological.
- The Einstein-Hilbert, Palatini and Holst actions all agree on-shell.
  - Using only “half” the equations of motion  $T_I = 0$ .
  - The bulk and boundary terms agree separately.

# The imaginary part of the GR action



In GR:

$$\text{Im } S = \frac{1}{16G} \sum_{\text{flips}} A_{\text{flip}}$$

( $A$  - area)

In more general theories:

$$\text{Im } S = \frac{1}{4} \sum_{\text{flips}} \sigma_{\text{flip}}$$

( $\sigma$  - Wald entropy functional)

# What do you mean, an imaginary part?

## This sounds ridiculous

Lorentzian actions are supposed to be real.

- Unitary amplitudes  $e^{iS}$ .

Euclidean actions are imaginary.

- Exponentially damped amplitudes  $e^{-\text{Im } S}$ .
- Tunneling, thermal behavior.

## But in gravity...

- Classical Lorentzian solutions can have thermal properties.
  - Black holes, other horizons.
- Unitarity is not at all understood for general spacetime regions.
  - Where it *is* understood ( $S$ -matrix),  $\text{Im } S$  vanishes.

## Deriving $\text{Im } S$

Arises from the action's boundary term, integrated over the infinitesimal neighborhood  $\delta\mathcal{F}$  of a flip surface  $\mathcal{F}$ :

$$\text{Im } S = \frac{1}{8\pi G} \sum_{\text{flips}} \text{Im} \int_{\delta\mathcal{F}} \sqrt{\frac{-h}{n \cdot n}} K d^3x$$

In first-order formulation - arises from the  $n_I dn_J$  piece of the boundary term:

$$\text{Im } S = \frac{1}{8\pi G} \sum_{\text{flips}} \text{Im} \int_{\delta\mathcal{F}} \Sigma^{IJ} \wedge \frac{n_I dn_J}{n \cdot n}$$



## Simplest example

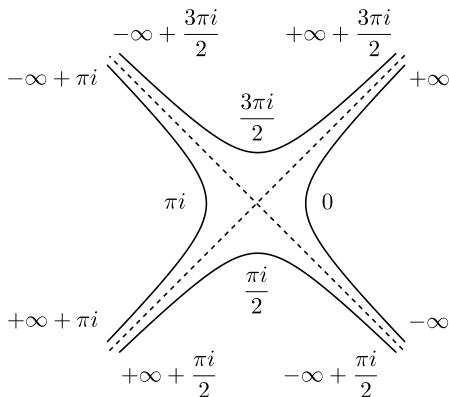
- 1+1d:  $\Sigma^{IJ} \rightarrow \epsilon^{IJ}$
- Flat space, trivial gauge:  $A_{IJ} = 0$

$$S = \frac{1}{8\pi G} \int \sqrt{\frac{-h}{n \cdot n}} K dx = \frac{1}{8\pi G} \int \epsilon^{IJ} \frac{n_I dn_J}{n \cdot n} = \frac{1}{8\pi G} \int d\alpha$$

- In Euclidean:  $\int d\alpha = 2\pi$
- In Lorentzian:  $\int d\alpha = 2\pi i$  [Sorkin, 1974]

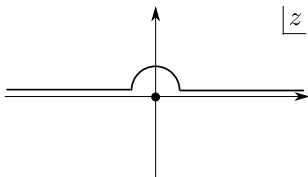
## Angles in the Lorentzian plane

When  $n_l$  flips signature (crosses a null direction), its angle jumps by  $\pi i/2$ .



## Derivation of the $\pi i/2$ jump:

- Write your vector as:  $n^\mu \sim L^\mu + z\ell^\mu$ .
- Boost angle is:  $d\alpha = \frac{dz}{2z}$ .
- At signature flip,  $z$  crosses through zero.
- Must deform the integration contour to avoid the pole!



$$\text{Im } \alpha = \frac{1}{2i} \oint \frac{dz}{2z} = \frac{\pi}{2}$$

## Back to GR in 1+1d

### In flat space:

- The action for a region in flat space is:  $S = \frac{1}{8\pi G} \cdot 2\pi i = \frac{i}{4G}$ .
- Arises from the 4 poles in the  $\frac{n_I dn_J}{n \cdot n}$  boundary term at signature flips (null  $n_I$ ).
- Sign chosen so that  $\text{Im } S > 0$ .  
 $\Rightarrow$  Amplitudes  $e^{iS} \sim e^{-\text{Im } S}$  not exploding.

### In curved space (in presence of matter):

- Only the real part changes. Still have:  $\text{Im } S = \frac{1}{4G}$ .
- The  $\frac{n_I dn_J}{n \cdot n}$  integral with its poles remains unchanged.
- The bulk term and the  $A_{IJ}$  boundary term have no poles.

## Upgrade to 3+1d (or any other dimension)

$$\text{Im } S = \frac{1}{8\pi G} \sum_{\text{flips}} \frac{\pi}{2} \cdot A_{\text{flip}} = \frac{1}{16G} \sum_{\text{flips}} A_{\text{flip}}$$

- $\text{Im } S$  is due to poles in the  $\Sigma^{IJ} \wedge \frac{n_I dn_J}{n \cdot n}$  boundary term.
- Poles now appear at 2d signature-flip surfaces (null  $n_I$ ).

On a flip surface, only one component contributes to the numerator

- Form indices on  $\Sigma^{IJ}$  - along the flip surface.
- Form index on  $dn_J$  - along the null normal/tangent.
- $IJ$  indices - in the 1+1d plane transverse to the flip surface.

The calculation reduces to 1+1d

- The component of  $\Sigma^{IJ}$  provides an area-density factor.

# Classical action of a flat 4-simplex

$$S = \sum_{\text{2d corners}} \frac{A\theta}{8\pi G}$$

$$\text{Im } S = \sum_{\text{thin corners}} \frac{A}{8G}$$

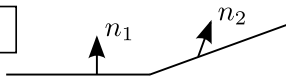
## 4-simplex with spacelike 3d faces in Minkowski space

- $S$  is purely a boundary term, concentrated at 2d corners (triangles).
- Dihedral angles  $\theta$  - discrete version of extrinsic curvature.

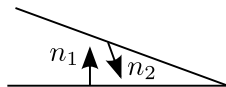
## Angles $\theta$ are complex

- Every “thin” corner hides two signature flips.

$$\text{Im } \theta = 0$$



$$\text{Im } \theta = \pi$$



# EPRL spinfoam amplitude for a large- $j$ 4-simplex

[Barrett, Dowdall, Fairbairn, Hellmann & Pereira, 2010]

$$\mathcal{A} = N_+ e^{iS_+} + N_- e^{iS_-}$$

$$S_{\pm} = \sum_{\text{corners}} j(\mp \gamma \operatorname{Re} \theta - \operatorname{Im} \theta)$$

$\pm$  denotes the two stationary points of the “path integral”.

- Correspond to two parity-related configurations of the 4-simplex.
- Can isolate one of the two by smearing over spins  $j$ . [Bianchi & Ding, 2011]

Expression for  $S_{\pm}$  adapted from Barrett et.al. by comparing the angle conventions.

- $\operatorname{Re} \theta$  is  $-\Theta$  from Barrett et.al.
- $\operatorname{Im} \theta$  is  $\Pi = 0, \pi$  from Barrett et.al.

## From spins to areas

$$A = 8\pi G \sqrt{\gamma^2} j$$

$$S_{\pm} = \sum_{\text{corners}} \frac{\text{sign}(\gamma) A}{8\pi G} \left( \mp \text{Re } \theta - \frac{1}{\gamma} \text{Im } \theta \right)$$

In the area spectrum, used  $\sqrt{\gamma^2} = \text{sign}(\gamma) \gamma$

- Rather than  $\gamma$  - to cover both signs of  $\gamma$ .
- Rather than  $|\gamma|$  - to be as complex-analytic as possible.

Two halves of the complex  $\gamma$  plane

- $\sqrt{\gamma^2}$  and  $\text{sign}(\gamma)$  can be analytically continued from the real line.
- However, only in two disjoint domains of the complex plane.
- Positive and negative real  $\gamma$  lie in different domains.
- Away from the real line, the domain boundary is arbitrary.



## Comparing to the classical action

$$S_{\pm} = \sum_{\text{corners}} \frac{\text{sign}(\gamma)A}{8\pi G} \left( \mp \text{Re } \theta - \frac{1}{\gamma} \text{Im } \theta \right)$$

### Old interpretation

- For real  $\gamma$ , the second term is an integer multiple of  $\pi \Rightarrow$  throw away.
- The first term reproduces the classical action, up to sign.

### New interpretation

- The first term reproduces only the real part of the classical action.
- The imaginary part is reproduced by the second term, iff analytically continue  $\gamma \rightarrow \text{sign}(\gamma)i$ .
  - Sign chosen so that  $\text{Im } S_{\pm} > 0$ .
  - $\gamma > 0$  goes to  $\gamma = i$ ;  $\gamma < 0$  goes to  $\gamma = -i$ .
  - This is allowed - we had two separate analytic domains to begin with.

Let's send  $\gamma$  to  $\text{sign}(\gamma)i$

$$S_{\pm} \rightarrow \mp \text{sign}(\gamma) \text{Re } S_{\text{class}} + i \text{Im } S_{\text{class}}$$

### Properties of the analytically-continued amplitude

- $S_{\pm}$  reproduces the classical action, up to the sign of the real part.
- $e^{iS_+}$  and  $e^{iS_-}$  are still complex-conjugate to each other.

### What does $\gamma = \pm i$ mean?

- Back to the self-dual Ashtekar connection!
- However, don't know how to quantize with non-real  $\gamma$ .
- Use real  $\gamma$  as a regulator - remove at the end of the calculation.

# Black hole state-counting in LQG

Does the entropy from state counting agree with the Bekenstein-Hawking formula?

- Bekenstein-Hawking is derived from a classical action.
- To compare, must express the state-counting entropy through parameters of an effective action in a classical limit.
- In LQG, there is no calculation of a continuum effective action to compare with.

The large-spin limit is a semiclassical regime where we have an action to compare with!

- Carry out the LQG state counting, with restriction to large spins and fixed number of punctures. [Frodden, Geiller, Noui & Perez, 2012]
- Bekenstein-Hawking reproduced, iff  $\gamma$  is analytically continued to  $\pm i$ .
- $\gamma = \pm i$  leads to a correct action *and* a matching entropy!

## Large-spin limit vs. the continuum

Large-spin LQG with  $\gamma \rightarrow \pm i$  reproduces key aspects of continuum GR.

- However, this happens non-trivially.

In the continuum:

- $\gamma$  doesn't matter.
- $\text{Im } S$  comes from the Palatini boundary term.

In the large-spin limit:

- $\gamma$  matters.
- $\text{Im } S$  comes from a Holst-type  $1/\gamma$  term. Large quantum effect!

The two limits seem distinct.

- How should we understand this?

# An RG flow between large spins and the continuum?

## A tempting interpretation

$\gamma^2$  “flows” from arbitrary values in the continuum to the self-dual value at large spins.

## A suggestive coincidence

- In a 1-loop perturbative calculation (in Euclidean, with fermions),  $\gamma^2$  flows from arbitrary values in the IR to the self-dual value in the UV.

[Benedetti & Speziale, 2011]

- Sounds familiar...
- The continuum is certainly “in the IR”.
- Is the large-spin limit somehow “in the UV”?

# The large-spin limit as a high-energy regime

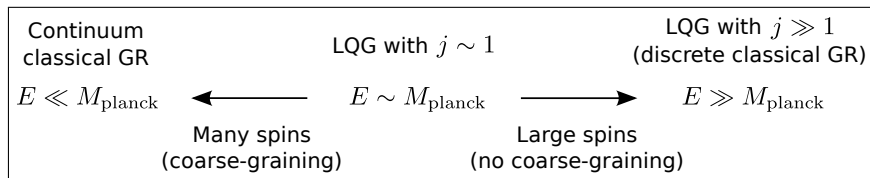
## UV and IR are misleading words

- The large-spin limit describes large *distances*  $L \sim \sqrt{j} \ell_{\text{planck}}$ .  
In quantum gravity, this can mean either large or small *energies*  $E$ .
- In QFT: small energies  $E \sim \hbar/L = (\ell_{\text{planck}}/L)M_{\text{planck}}$
- In classical GR: large energies  $E \sim L/G = (L/\ell_{\text{planck}})M_{\text{planck}}$

## What's the appropriate physical picture?

- QFT picture:
  - Long-wavelength gravitons *perturbing* a background.
  - Action (of graviton) over distance  $L$  is  $S/\hbar \sim 1$ .
- Classical GR picture:
  - Large black holes *defining* a background.
  - Action (of classical metric) over distance  $L$  is  $S/\hbar \sim (L/\ell_{\text{planck}})^2 \gg 1$ .
- The large-spin limit is more like the classical GR picture.
- Should be thought of as high-energy:  $E/M_{\text{planck}} \sim \sqrt{j} \gg 1$ .

# Two classical limits of LQG



- Continuum limit: many low-energy quanta.
- Large-spin limit: few high-energy quanta.

- All this refers to the large-spin limit *of the fundamental theory*.
- Large spins may also arise in a coarse-graining picture. This will involve some new, effective dynamics.

# Summary (again)

## Classical GR

- First-order and second-order actions agree on-shell.
- The action has a nonvanishing imaginary part:

$$\text{Im } S = \frac{1}{16G} \sum_{\text{flips}} A_{\text{flip}} = \frac{1}{4} \sum_{\text{flips}} \sigma_{\text{flip}}$$

## The large-spin limit of LQG

- The large-spin vertex amplitude reproduces the classical action iff analytically continued to  $\gamma = \pm i$ .
- A first-ever agreement in LQG between BH entropy derived from an effective action and from a microscopic state counting.
- The limit should be thought of as a high-energy (“transplanckian”) regime.



# New questions

## Understand $\text{Im } S$ !

- Proper relation to black hole entropy?
- Implications on unitarity?

## The large-spin limit wants $\gamma = \pm i$ .

- Must be a way to quantize with  $\gamma = \pm i$  directly!
- Use lightlike structures?

## New understanding of large-spin limit?

- Conceptual implications on “IR” divergences?