

SIMPLICITY CONSTRAINTS AND THE IMMIRZI PARAMETER IN DISCRETE QUANTUM GRAVITY

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Summary

introduction and brief review

set up of discrete canonical Plebanski

reduction, part 1: simplicity constraints

reduction, part 2: gluing constraints

conclusions and outlook

Introduction and brief review

Plebanski action with Holst term:

$$\mathcal{S}_{Pleb, \mathcal{M}}[X, w, \phi] = \frac{1}{2} \int_{\mathcal{M}} \left[\text{tr}(X \wedge F[w]) + \frac{1}{\gamma} \text{tr}(\star X \wedge F[w]) + \frac{1}{2} \phi_{AB} \star X^A \wedge X^B \right]$$

SO(4) BF term

↑
Barbero-Immirzi parameter

Holst term

Simplicity constraints

some definitions: $\text{tr}(X_1 X_2) = X_1^A \delta_{AB} X_2^B$

$$\star X^A = \epsilon^{AB} X^B$$

Introduction and brief review

Lagrange multiplier:

$$\phi^{AB} = \phi^{BA}$$

symmetric

$$\epsilon_{AB} \phi^{AB} = 0$$

traceless

Simplicity constraints:

$$\star X^A \wedge X^B = \frac{V}{4!} \epsilon^{AB}$$

where

$$V = \star X^A \wedge X^B \epsilon_{AB} = \text{tr}(X \wedge X)$$

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Non-degenerate solutions:

assuming non-degeneracy, that is: $V \neq 0$

$$X = \begin{cases} \pm(e \wedge e) \\ \pm \star(e \wedge e) \end{cases}$$

topological sector
gravitational sector

sectors become mixed

without Holst term

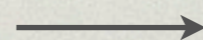
with Holst term

choosing the second sector of solutions

$$\mathcal{S}[e, w] = \frac{1}{2} \int_{\mathcal{M}} \left(\text{tr}((e \wedge e) \wedge F[w]) + \frac{1}{\gamma} \text{tr}(\star(e \wedge e) \wedge F[w]) \right)$$

solving half the equations of motion

trivializes the Holst term



no effect on classical dynamics

Introduction and brief review

Spinfoam method:

consider a triangulated manifold

$$\mathcal{M} \longrightarrow \Delta_4$$

$$X_{\mu\nu}{}^A \longrightarrow X_f$$

$$w_\mu{}^A \longrightarrow M_t = \mathcal{P} \exp \left(\int w \right)$$

start with a quantum theory of *BF* + *Holst* terms only

$$\mathcal{Z}_\Delta = \int \mu_\Delta[X_f, M_t] e^{i\mathcal{S}_\Delta[X_f, M_t]}$$

$$\mathcal{S}_\Delta[X_f, M_t] = \sum_{f \in \Delta} X_f M_f$$

advantage: well-defined measure

Introduction and brief review

Spinfoam method continued:

integrate bivectors explicitly

$$\mathcal{Z}_\Delta = \int \prod_{t \in \Delta} dM_t \delta(M_f)$$

re-expand in terms of representations of $\text{SO}(4) \sim \text{SU}(2) \otimes \text{SU}(2)$

magnetic basis

$$|j_+, m_+ \rangle \otimes |j_-, m_- \rangle$$

coherent state basis

$$|j_+, \vec{n}_+ \rangle \otimes |j_-, \vec{n}_- \rangle$$

quantum bi-vector

reason: amplitude in localized chunks associated to tetrahedra and 4-simplices

making contact with boundary formalism

allows the imposition of simplicity constraints as operator constraints

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Spinfoam method continued:

impose simplicity constraints

remember: $\star X^A \wedge X^B = \frac{V}{4!} \epsilon^{AB}$

on the 4-simplices these split up as

diagonal simplicity	simplicity of bi-vector on each triangle
cross simplicity	pairwise simplicity for triangles sharing an edge
volume simplicity	pairwise simplicity for triangles sharing a vertex

turns out that we do not need volume simplicity

diagonal + cross on **BF amplitude** implies **volume**

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Spinfoam method continued:

Result

EPRL/FK spin foam models

[Engle, Pereira, Rovelli '08]

[Engle, Livine, Pereira, Rovelli '08]

[Freidel, Krasnov '08]

[Livine, Speziale '08]

wonderful property

boundary state space coincides with that of LQG restricted to a single graph

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Take-away point:

Immirzi parameter plays a non-trivial role

in both

spin foam models

and

loop models

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Some remaining questions:

spin foam quantization is not a conventional method of quantization

————→ check that it gives gravity in the semi-classical limit

looking for gravitons, cosmological sectors etc.

[Rovelli '06 and subsequently]

[Bianchi, Rovelli, Vidotto '10]

[Ashtekar, Campiglia, Henderson '09]

————→ check correspondence to Dirac quantization

