

Panel on Boundary Modes and Celestial Holography

Wolfgang Wieland

Austrian Academy of Sciences
Institute for Quantum Optics and Quantum Information
University of Vienna

www.wmwieland.eu

ILQGS

03-05-2022

Corners break symmetries and turn otherwise unphysical gauge directions into physical boundary modes.

Origin of intriguing developments in our field.



At the corner, symmetries are broken: corner conflict in classical architecture.

- complementary representation of LQG quanta of area.
- quasi-local observables in no-perturbative quantum gravity.
- gluing and coarse graining.
- towards radiative modes and scattering in LQG.
- Wheeler De Witt equation on null surface = Ward identities.
- Quantum reference frames.
- BH entropy/area counting.

Pre-symplectic potential for first-order gravity*

$$\Theta_{\Sigma} = \frac{1}{16\pi G} \left[\int_{\Sigma} *(e_{\alpha} \wedge e_{\beta}) \wedge \mathbb{d}A^{\alpha\beta} + 2 \oint_{\partial\Sigma} *(e_{\alpha} \wedge e_{\beta}) n^{\alpha} \mathbb{d}n^{\beta} \right].$$

Are the radiative modes the only Dirac observables that we have?

To know the dimension of a manifold it is enough to know the dimension of its tangent space.

Warm up: Take **linearised gravity** in first-order variables

$$\begin{aligned} e^{\alpha} &= \Lambda^{\alpha}_{\mu} (dX^{\mu} + f^{\mu}), \\ A^{\alpha}_{\beta} &= \Lambda^{\alpha}_{\mu} d\Lambda_{\beta}^{\mu} + \Lambda^{\alpha}_{\mu} \Delta^{\mu}_{\nu} \Lambda_{\beta}^{\nu}. \end{aligned}$$

Evaluate Ω_{Σ} in a round ball centred at origin in X^{μ} -space.

* see [Corichi, Wilson-Ewing 2011; ww 2011, ww 2017; Freidel, Geiller, Pranzetti 2020; Bodendorfer, Neiman 2013; Margalef-Bentabol, Villaseñor, Barbero G. 2021; ...]

Only at the corner, do X^μ -fluctuations become physical

$$\begin{aligned} \Omega_\Sigma \Big|_{Minkowski} &= \frac{1}{8\pi G} \int_\Sigma dX_{[\mu} \wedge \mathbb{d}f_{\nu]} \wedge \mathbb{d}(*\Delta^{\mu\nu}) + \\ &\quad - \frac{1}{8\pi G} \oint_{\partial\Sigma} \left[\underbrace{\mathbb{d}(*F_{\mu\nu}) \wedge X^{[\mu} \mathbb{d}X^{\nu]}}_{\rightarrow \text{ADM charges at spi}^*} + \right. \\ &\quad \quad + d^2 v \rho \varepsilon^{ab} \varepsilon^{cd} \mathbb{d}K_{ac} \wedge D_{(b} \mathbb{X}_{d)}^\downarrow + \\ &\quad \quad + d^2 v \rho \varepsilon^{ab} \varepsilon^{cd} \mathfrak{g}_{ac} \wedge D_b D_d \mathbb{T}, \\ &\quad \quad \left. - \frac{1}{2} d^2 v^a (\mathfrak{g}_{ab} - h_{ab} \mathfrak{g}_d^d) \wedge \mathbb{N}^b \right]. \end{aligned}$$

3+1 split of basic variations

$$\partial_\mu^a \mathbb{d}X^\mu \equiv \mathbb{X}^a = \mathbb{T}n^a + \mathbb{X}_\downarrow^a.$$

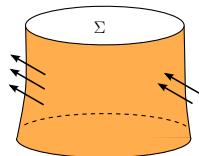
* [Ashtekar, Hansen 1978]

In linearised gravity, diffeo charges are trivially integrable (as in 2+1). Not so in full non-linear theory (**because of radiation**). No canonical generators for time-like (or radial) diffeos. **Subsystems characterised by charges, flux of radiation and choice of boundary embedding.**

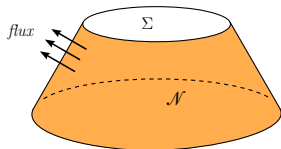
- Choice for how to extend the corner $\partial\Sigma$ into a worldtube \mathcal{N} .
- Choice for what is the flux of radiative modes crossing the boundary.
- Most ingenious idea [Freidel, Barnich, Troessaert, Leigh, Ciambelli], see also [Riello, Gomes], to study Komar charges on phase space via field-space connection. More investigation needed.

$$\Omega_{\Sigma}^{new} = \mathbb{D}\Theta_{\Sigma}^{new}, \quad \mathbb{D} = \mathbb{d} - \mathcal{L}_{\mathbb{X}},$$

$$\Omega_{\Sigma}^{new}(\delta, \mathcal{L}_{\xi}) = \mathbb{D}_{\delta}[\Theta_{\Sigma}^{new}(\mathcal{L}_{\xi})] - \underbrace{\Theta_{\Sigma}^{new}(\mathcal{L}_{[\mathbb{X}(\delta), \xi]})}_{flux}.$$



vs.



Bulk plus boundary action:

$$S = \frac{i}{8\pi\gamma G}(\gamma + i) \left[\int_{\text{bulk}} \Sigma_{AB} \wedge F^{AB} + \int_{\text{null-boundary}} \eta_A \wedge (D - \frac{1}{2}\varkappa)\ell^A \right] + \text{cc.}$$

Boundary conditions along \mathcal{N} : $\delta[\varkappa_a, l^a, m_a]/\sim = 0$.

Covariant pre-symplectic potential for the partial Cauchy surfaces:

$$\Theta_\Sigma = \frac{i}{8\pi\gamma G}(\gamma + i) \left[\int_{\text{disk}} \Sigma_{AB} \wedge \mathbb{d}A^{AB} - \oint_{\text{corner}} \eta_A \mathbb{d}\ell^A \right] + \text{cc.}$$

Heisenberg algebra at the two-dimensional corner

$$\{\eta_{Aab}(z), \ell^B(z')\}_{\mathcal{C}} = -\frac{8\pi i \gamma G}{\gamma + i} \delta_A^B \tilde{\delta}^{(2)}(z, z') \underline{\xi}_{ab}.$$

$$\widehat{\text{Area-flux}}[\mathcal{C}] \Psi_{\text{phys}} = 4\pi\gamma\hbar G/c^3 \oint_{\mathcal{C}} [a_A^\dagger a^A - b_A^\dagger b^A] \Psi_{\text{phys}}.$$

Signature (0++) metric.

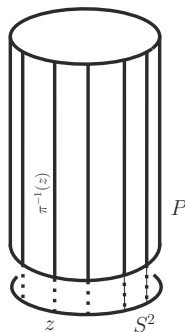
$$q_{ab} = \delta_{ij} e^i_a e^j_b, \quad i, j = 1, 2.$$

Parametrisation of the dyad

$$e^i = \Omega S^i_j e^j_{(o)}.$$

Choice of time:

$$\partial_U^b \nabla_b \partial_U^a = -\frac{1}{2} \left(\Omega^{-2} \frac{d}{dU} \Omega^2 \right) \partial_U^a$$



Kinematical phase space for radiation: $\mathcal{P}_{kin} = \mathcal{P}_{abelian} \times T^*SL(2, \mathbb{R})$.

$$\Theta_{\mathcal{N}} = \frac{1}{8\pi G} \int_{\mathcal{N}} d^2v_o \wedge \left[p_K d\tilde{K} + \frac{1}{\gamma} \Omega^2 d\tilde{\Phi} + \tilde{\Pi}^i_j [S dS^{-1}]^j_i \right] + \text{corner term.}$$

Abelian variables:

$U(1)$ connection: $\tilde{\Phi}$, area: $\Omega^2 d^2v_o$, lapse: $\tilde{K} := dU$, expansion: p_K .

Upon imposing 2nd-class constraints: Dirac bracket for radiative modes

$$\{S^i_m(x), S^j_n(y)\}^* = -4\pi G \Theta(U_x, U_y) \delta^{(2)}(\vec{x}, \vec{y}) \Omega^{-1}(x) \Omega^{-1}(y) \\ \times \left[e^{-2i(\Delta(x) - \Delta(y))} [XS(x)]^i_m [\bar{X}S(y)]^j_n + \text{cc.} \right].$$

Gauge symmetries:

- 1 $U(1)$ transformations
- 2 vertical diffeomorphisms along null generators

Main results discussed:

- Boundary conditions for radiative data altered by Barbero–Immirzi parameter.
- Barbero–Immirzi parameter mixes $U(1)$ frame rotations and dilations. **This is an important observation – it is the geometric origin for LQG quantum discreteness of area.**
- Poisson brackets for the boundary modes altered by addition of the Immirzi parameter. Poisson brackets for radiative modes unchanged.
- New representation of quantum geometry.

Future of the programme:

- Ward identities/constraints linking corner data to flux.
- State-amplitude correspondence as we know it from spinfoams: physical states: $\Psi_{\mathcal{N}}(\text{corner data}_+ | \text{radiative data} | \text{corner data}_-)$.

$$\langle \text{out} | \text{in} \rangle_{\text{EPRL-foam?}} = \sum_C \bar{\Psi}_{\mathcal{N}^-} (c_{+\infty} | \text{out} | C) \Psi_{\mathcal{N}^+} (C | \text{in} | c_{-\infty}).$$

- Impulsive waves and spinfoams, building upon [ww2016].